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# The DO-climate events are noise induced: statistical investigation of the claimed 1470 years cycle

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Abstract

The significance of the apparent 1470 years cycle in the recurrence of the Dansgaard-Oeschger (DO) events, observed in the Greenland ice cores, is debated. Here we present statistical significance tests of this periodicity. The detection of a periodicity relies strongly on the accuracy of the dating of the DO events. Here we use both the new NGRIP GICC05 time scale based on multi-parameter annual layer counting and the GISP2 time scale where the periodicity is most pronounced. For the NGRIP dating the recurrence times are indistinguishable from a random occurrence. This is also the case for the GISP2 dating, except in the case where the DO9 event is omitted from the record. Whether or not the record shows a truly periodic beating has strong implications for identifying the underlying cause. If the recurrence is periodic it suggests an external cause. If the recurrence of DO events is not periodic it points to triggering mechanisms internal to the climate system being manifested at the millennial timescale.

1 Introduction

The 1470 years period was first noted as a significant peak in the spectral density of the GISP2 isotope record, which is based on annual layer counting (Grootes and Stuiver, 1997). However, the peak was not significant using the GRIP based “ss09sea-model” timescale (Johnsen et al., 2001; Ditlevsen et al., 2005). The new stratigraphic NGRIP (North GRIP members, 2004) GICC05 time scale is based on multi parameter annual layer counting (Andersen et al., 2006). The timescale is similar to the GISP2 stratigraphy over longer time periods but very different across the DO events, as annual accumulation rates are more closely linked to climate in GICC05. This means that the warm interstadials generally have a longer (probably erroneous) duration using the GISP2 time scale than they do using the NGRIP time scale (Svensson et al., 2006). In this analysis we use both the GISP2 (Meese/Sowers) and the NGRIP (GICC05) time

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scales. Furthermore, we shall limit the test to the period 11–42 kyr BP, the current limit of GICC05, where the signature of the 1470 year cycle is most pronounced and the dating highly reliable. The preliminary extension of GICC05 shows substantial differences to the GISP2 dating before 40 kyr BP. This may indicate that the GISP2 record prior to 40 kyr BP becomes increasingly less trustworthy for detecting a periodic signal. Even though we perform our analysis for both the GISP2 and the NGRIP dating we will emphasize that there are good reasons to believe that the new multi-parameter annual layer counting applied to the NGRIP core is the most reliable and accurate of the two (Svensson et al., 2006).

## 2 The 1470 years period

The DO events have a characteristic saw-tooth shape, beginning with a very abrupt transition from the glacial (stadial) state into the DO (interstadial) state. This is followed by a gradual decrease in the  $\delta^{18}\text{O}$  isotope ratio until eventually there is a smaller jump back into the stadial state. This strongly non-sinusoidal shape of the climate curve results in a large part of the spectral power in a signal being spilled into overtones, which can result in a lowering of a spectral peak below the noise level. Thus there is a tendency of underestimating the significance of periodical components in a spectral analysis of such a signal. By bandpass filtering the 1470 years spectral component can be removed all together, while this does not remove the DO events from the record (Wunsch, 2000). Furthermore, if the periodicity is such that every now and then a periodic jump is “skipped”, as is the case in a stochastic resonance (Alley et al., 2001), this will also reduce the power in a spectral peak. Therefore it is advantageous to focus on the timing of the well-defined abrupt jumps into the interstadial states. By doing that it was noted previously that these initiations line up well with a constant beating of 1470 years (Schulz, 2002). These times are shown as vertical markers in Fig. 1. The two curves are the isotope records from NGRIP (red) and GISP2 (blue) on the two different time scales.

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The apparent regular timing suggests a periodic forcing such as an hitherto undiscovered solar period, or a beating of several periodic forcings (Braun et al., 2005). This has, however, recently been refuted in a comparison between the  $^{10}\text{Be}$  and the  $\delta^{18}\text{O}$  records from the GRIP icecore (Muscheler and Beer, 2006). The regular timing is quite striking but needs to be tested statistically. This is not completely straight forward. The general problem is that when observing a pattern in a data set, the significance of the pattern can be very difficult to assess a posterior unless the space of possible outcomes for “striking patterns” is known.

### 3 Defining DO events

The starting point for the analysis is to decide on criteria for defining DO events and determining the transition times. This has previously been done in a variety of ways: The “canonical” numbered DO were identified visually (Dansgaard et al., 1993), Schulz defined the DO events from a positive 2 permil anomaly in the 12 kyr high-pass filtered isotope signal. By that DO9 is disregarded. Rahmstorf defines a criterion of increase of 2 permil within 200 years on the 2-m sampled record (approx. 100 years low-pass). In this way DO9 is omitted and an event “A” in the Allerød period is included (Rahmstorf, 2003). Alley et al. use a bandpass procedure by which 43 events in the glacial period are defined (Alley et al., 2001). Ditlevsen et al. defined first upcrossings of an upper level following upcrossings of a lower level as criterion. In this way the critical dependence on the (arbitrary) low-pass filter and crossing levels is to a large extent avoided (Ditlevsen et al., 2005). Using this criterion several additional DO events are identified, such as DO2 which is split into two separate events.

Discussions of the criteria for defining the DO events will be deferred to a future publication. Here we simply apply our analysis to the different proposed DO event series. The absolute (cumulative) dating uncertainty for NGRIP (GICC05) is of the order 800 years at 40 kyr BP, while the uncertainty in the recurrence times is of the order 50 years. (Thus the last digit in the dating is insignificant) (Andersen et al., 2006).

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The reported dating uncertainty for GISP2 is approximately 1% down to 58 kyr BP, corresponding to approximately 20–50 years for the recurrence times (Meese et al., 1997). We expect this estimate to be somewhat optimistic (Svensson et al., 2006).

## 4 Measures of periodicity

5 We shall denote the identified time sequence for jumps as  $t_i, i=1, \dots, N$ . A preferred periodicity in the time sequence can be detected by the Rayleigh's R measure defined as  $R(\tau)=(1/N)|\sum_j \cos 2\pi t_j/\tau + i \sin 2\pi t_j/\tau|$ , where obviously  $R(\tau) \in (0, 1)$  (Huybers and Wunsch, 2005). This measure is easy to understand if we define the angles  $\theta_i=2\pi t_i/\tau$  and plot the angles on the unit circle. If the time sequence is multiples of the time  $\tau$  modulo an (unknown) phase, all angles will be located near the same point on the unit circle and  $R(\tau)\approx 1$ . On the contrary if the data points do not cluster on the unit circle we have  $R(\tau)\approx 0$ .

15 A second measure of the periodicity is the “Standard deviation of residuals” (Std. dev. res.). The residuals are defined as the distances of the data points from the (nearest) location of a perfect periodic signal. The phase and period of the periodic signal is chosen by optimization (Schulz, 2002). The measures were calculated for 5 cases, [1]: DO 0–10, NGRIP timescale (NG), [2]: DO 0–10, GISP2 timescale (G2), [3]: DO 0,A,1–8,10, NGRIP timescale (NG-DO9), [4]: DO 0,A,1–8,10–12, GISP2 timescale (G2-D09) (Rahmstorf, 2003), [5]: DO1c,1,2a,2b,3–10, NGRIP timescale (DKA-2005) (Ditlevsen et al., 2005). DO0 refers to the transition into the pre-boreal, while “A” is the Allerød event.

25 In Fig. 2, right panel, the value of  $R(\tau)$  as a function of  $\tau$  is shown for the two cases NG and G2. The period of 1470 years shows the largest value  $R=0.65$  ( $R=0.72$  for G2). The angles with respect to the 1470 years period of the time sequence of DO-jumps are plotted on the unit circle in Fig. 2, left panel. The mean phase is indicated by the radial line segments, the length is equal to  $R(1470 \text{ years})$ . The mean phase defines the vertical lines plotted in Fig. 1 (for the NGRIP time scale).

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The Rayleigh  $R$  and the Std. dev. res. for the records are listed in Table 1. Omitting DO9 as proposed by Rahmstorf [Rahmstorf \(2003\)](#) makes a big difference for the GISP2 dating, but not for the NGRIP dating.

## 5 Significance of period

5 The next, and necessary, step in the analysis is to test the significance of the periodicity found in the data. This can only be done by assuming a test-model generating the data. Given such a model, we may choose any measure derived from the data,  $x_d$  to compare with the same measure derived from similar realizations of the test-model,  $x_m$ . The null-hypothesis is then that the data series is a specific realization of the model. It  
10 is important to note that a null hypothesis can only be rejected and not confirmed. That is, the value of the chosen measure for the data may well be within the high likelihood region for the model, but this does not prove that the data cannot be generated from another (competing) model with same high likelihood for the chosen measure. On the contrary, only if the measure for the data falls within a low likelihood region, say with  
15 probability-measure  $p \ll 1$ , the model can be rejected with probability  $1-p$ .

## 6 Model 1: Exponential distribution

The simplest possible model which can be chosen for the statistical test is that the DO-events occur randomly, without a memory, on the millennial time scale. This is described by an exponential distribution for the waiting times corresponding to a Poisson  
20 process. The mean waiting time can be assumed to be 2800 years. This is obtained as an estimate from the mean waiting times for 14 DO-events in the period 10–50 kyr. This is also the estimate obtained from the best fit to an exponential distribution of all DO events in the full glacial period ([Ditlevsen et al., 2005](#)).

To test the data against this model we use the two measures, the Rayleigh  $R$  and  
25 the standard deviation of residuals, obtained from the data. For each of the measures

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a probability density for a sample, similar to the observed record, is obtained from a Monte Carlo generated ensemble of 1000 realizations. The results are shown together with the measures from the data records in Fig. 3, top panels. To account for the dating uncertainty an additional uncorrelated gaussian noise, with  $\sigma=100$  yr corresponding to a conservative error estimate, was added to the model signal. In terms of the probability density,  $p(t)=\exp(-t/\tau)\Phi[(t-\sigma^2/\tau)/\sigma]/\tau$  substitutes the exponential density  $p(t)=\exp(-t/\tau)/\tau$ , where  $\sigma$  is the standard deviation of the dating noise and  $\Phi[x]$  is the error function. With  $\sigma<200$  yr this had an insignificant influence on the result and is for simplicity omitted in the following.

From the figure it is obvious that the data records fall within the high likelihood region of the exponential distribution for both measures. The 90% (dashed) and 99% (full) confidence levels are shown in the figure as vertical blue bars. Note that the confidence levels are accurately calculated from the cumulated distribution, independent of the binning used for the histogram. Thus there is no basis for rejecting the hypothesis of no-periodicity for the data. An exception is the curious case of the GISP2 dating with DO9 omitted, in which case the model can be rejected at the 99% confidence level.

## 7 Model 2: Periodic beating

The opposite proposition of rejecting a periodic component depends on the additional independent noise in the signal. Assuming a perfect periodic beating, with occasional omissions, blurred by the dating noise, the time series model to test data against is  $t_i=\tau_0+n(i)\tau$ , where  $\tau=1470$  yr,  $n(i)$  is a monotonous integer function and  $e_i$  is an independent unit variance gaussian noise. The standard deviation  $\sigma$  is taken to be 100 years corresponding to the conservative estimate for the dating of the two ice-core records. The results from a simulation of 1000 realizations of this model are shown in Fig. 3, middle panels. This model can be rejected at the 99% confidence level in all 5 cases. Now it is highly improbable that a hitherto undiscovered period of such dominance should exist in the climate system. The climate system is dominated by internal

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noise masking possible periodic components. Thus a more reasonable assumption is that a periodicity is caused by an internal non-linear amplification of a weak external periodic forcing. This could be described by a stochastic resonance as proposed by Alley et al. (2001).

### 8 Model 3: Stochastic resonance

The stochastic resonance model Benzi et al. (1982) is defined by the governing equation:

$$dx = -\partial_x U_a(x, t, \tau)dt + \sigma dB = \{-2(x^3 - x) + a \cos(2\pi t/\tau)\}dt + \sigma dB,$$

where a particularly simple form of the drift potential  $U_a(x, t, \tau)$  is chosen here. The potential is a double-well potential, which changes periodically with period  $\tau$  from having a shallow well (s) to the right and a deep well (d) to the left to the opposite situation. The ratio of the barrier heights  $H_{s \rightarrow d}/H_{d \rightarrow s}$  is determined by the model parameter  $a$ . The time scales for jumping from the shallow well to the deep well is given by an Arrhenius formula;  $T_{s \rightarrow d} \sim \exp(H_{s \rightarrow d}/2\sigma^2)$ , and similarly for  $T_{d \rightarrow s}$ . The criterion for resonance, where the signal  $x$  is most periodic, is  $T_{s \rightarrow d} \ll \tau \ll T_{d \rightarrow s}$ . This determines the noise intensity  $\sigma$ .

The proposition of rejecting a stochastic resonance (SR) model for the ice-core data is more tricky, since there exists a continuum of SR-models with waiting time distributions from the exponential to the delta-distribution for the perfect periodicity (Ditlevsen et al., 2005). However, the only spectral weight notably above the continuum is at  $\tau^{-1} = 1470 \text{ years}^{-1}$  (for the GISP2 dating) and not at the mean waiting time  $2800 \text{ years}^{-1}$ . Near the stochastic resonance one should expect the same order of magnitude “early jumps” (corresponding to a noise induced jump from the deep well to the shallow well) as “late jumps” (corresponding to missing a jump from the shallow to the deep well). The mean waiting time being about twice the observed spectral period indicates that a possible SR is “off the resonance” with a too low noise level. In terms of

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SR parameters, this means that the criterion;  $T_{s \rightarrow d} \ll \tau \ll T_{d \rightarrow s}$  ( $T_{s \rightarrow d}$  being the mean waiting time for a transition from the shallow to the deep well), is not fulfilled. We rather see  $\tau < T_{s \rightarrow d}$ .

Here we test against three SR-models with the period  $\tau=1470$  years and  $a=0.1, 0.2, 0.4$ . The mean number of DO-events being 11 events/31 kyr corresponding to the climate record. This determines the noise intensity to be  $\sigma=0.38, 0.35, 0.27$  for the three models.

An ensemble of 1000 simulations with same length as the data records were generated and the same three significance tests were performed. The results for the three models;  $a=0.1$  (light green),  $a=0.2$  (medium green)  $a=0.4$  (dark green) are shown in Fig. 3, panels (e) and (f). The distributions in panels (a) and (b), for the exponential distribution, are overplotted in gray in panels (e) and (f). It is seen that the first model,  $a=0.1$ , apparently has less periodicity, represented by Rayleigh's  $R$ , than the purely exponential model. This is because, in the case of the SR model, the distribution is of  $R(1470 \text{ yr})$ , while in the case of the exponential waiting time distribution (corresponding to  $a=0$ ), the distribution is for the largest value of  $R$  found in the sample. (Note that for the Nyquist frequency all points are aligned with  $R=1$ . This trivial limit is obviously excluded.) This means that for the SR model with  $a=0.1$ , the period will not be identified in comparison to other spurious coincidental periodicities. We have thus identified the "weakest" SR model which may be identified for a sample of the size of the record, and this SR model ( $a=0.2$ ) is less likely for the data than the exponential waiting time model.

## 9 Conclusions

In conclusion, the statistical tests show that the waiting times for DO events are within the high likelihood region of the exponential distribution (Figs. 3a, b). This distribution implies that there is no long term memory in the climate system or unknown 1470 years periodic forcing triggering the climate shifts. The assumption of the onsets being

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determined by a strictly periodic triggering (not activating at each period) masked by the dating uncertainty can with high significance be rejected (Figs. 3c, d) A remarkable exception for the rejection is the situation where DO9 is omitted for the GISP2 time scale. The relatively strong periodicity in that case is, however, not preserved in the newer NGRIP dating. By the nature of the statistical test we can only reject the hypothesis of a periodic component when the period is sufficiently above the noise level. For SR models with too low a strength of the periodic component, the period would with high probability not be detected in comparison to detecting a spurious coincidental periodicity in the sample (Figs. 3e, f).

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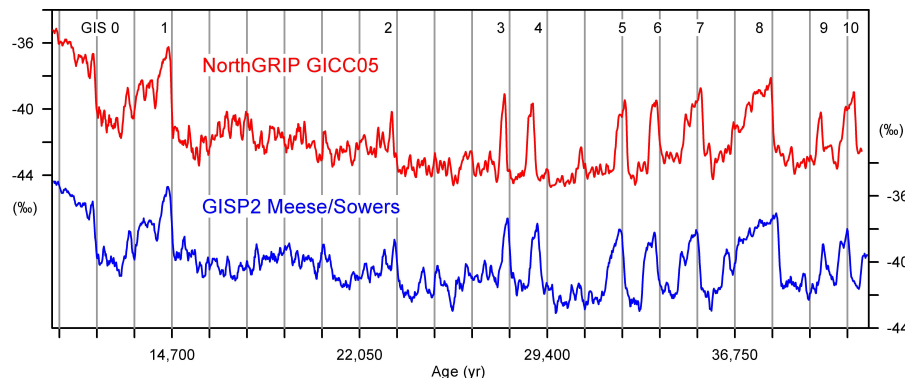
**Table 1.** The Rayleigh R and the Std. dev. of residuals for the 5 cases: NG: DO 0–10, NGRIP timescale, G2: DO 0–10, GISP2 timescale, NG-D09: DO 0,A,1–8,10, NGRIP timescale, G2-D09: DO 0,A,1–8,10–12, GISP2 timescale (Rahmstorf, 2003), DKA-2005: DO1c,1,2a,2b,3–10, NGRIP timescale (Ditlevsen, Kristensen and Andersen,2005). Note that the case G2-D09 is remarkably more periodic than the other 4 cases.

	Rayleigh R	Std. dev. res.
NG	0.65	0.92
G2	0.72	0.80
NG-D09	0.73	1.01
G2-D09	0.87	0.65
DKA-2005	0.60	0.94

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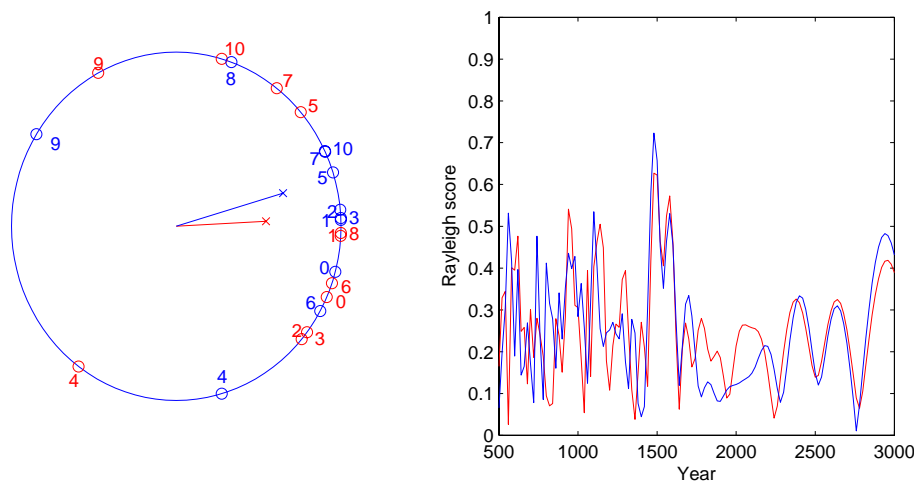


**Fig. 1.** The  $\delta^{18}\text{O}$  isotope records from NGRIP and GISP on their stratigraphic time scales (Alley et al., 1997; Andersen et al., 2006). The vertical bars are separated by 1470 years. The analysis focus on the well defined fast onsets of DO events, which are the transitions from the stadial to the interstadial states. Beginning at GIS0 the onset for the DO events are for the NGRIP GICC05 (GISP2) time scale: 11 700 (11 660); 13 130 (13 180); 14 680 (14 700); 23 340 (23 560); 27 780 (27 920); 28 900 (29 100); 32 500 (32 400); 33 740 (33 700); 35 480 (35 360); 38 220 (38 480); 40 160 (40 280); 41 460 (41 240). Ages are b2k=BP+50 years.

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**Fig. 2.** The Rayleigh R test for the two records. The maximum is obtained for the period  $\tau=1470$  years. Left panel shows the timing of the onsets  $t_n$  plotted on the unit circle using the transformation  $\theta_n=2\pi t_n/\tau$ . The red dots represents the NGRIP dating (NG) while the blue dots represents the GISP2 dating (G2). The segments of radians points at the mean phase, corresponding to the vertical bars in Fig. 1 (for NGRIP dating).

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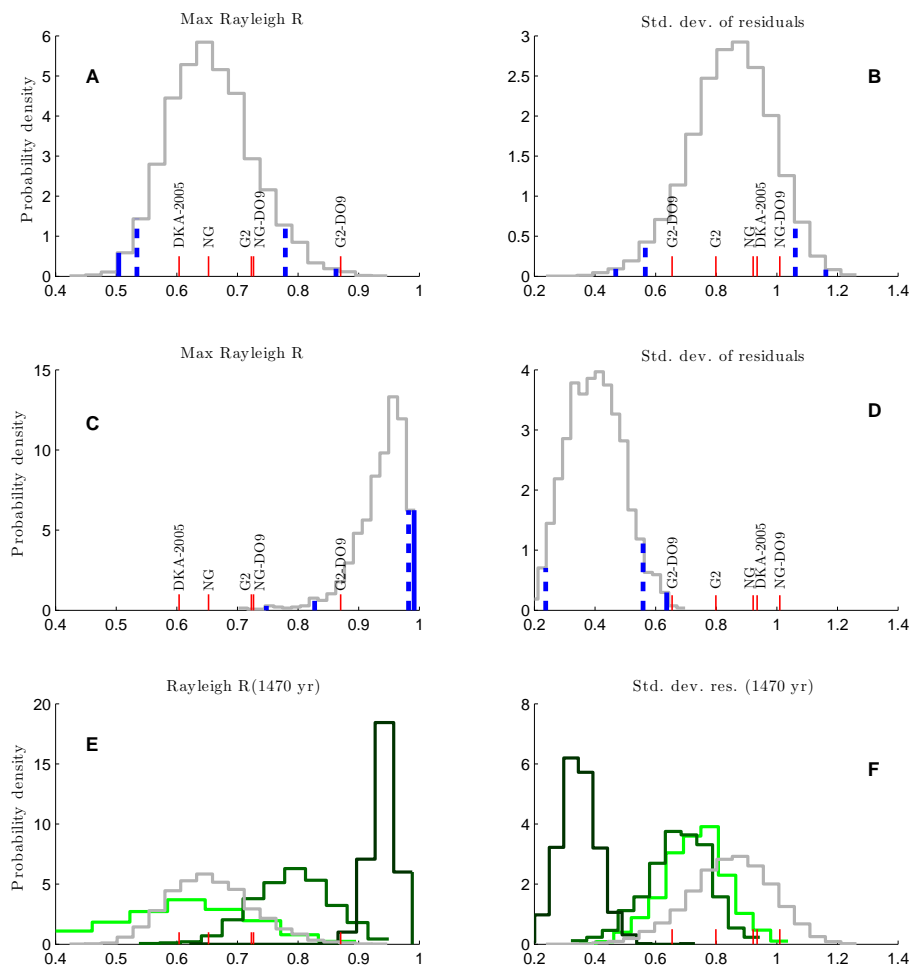


Fig. 3.

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**Fig. 3.** Panels **(A)** and **(B)**: By Monte Carlo an ensemble of 1000 realizations of waiting times in a 40 kyr period has been generated from an exponential distribution with mean waiting time of 2800 years, corresponding to 14 DO-events in 40 kyr. This gives probability densities for the maximal Rayleigh's  $R(\tau)$  in the range  $500\text{ yr} < \tau < 5000\text{ yr}$  and for the "Standard deviation of residual" (see text). The red bars give the values for the ice-core records (see text). The blue bars are 90% (dashed) and 99% (full) confidence levels. Panels **(C)** and **(D)**: Same as panels (A) and (B), where now the distribution functions are obtained for a perfect 1470 year periodic signal subject to a dating error taken to be a gaussian with standard deviation of 100 years. Panels **(E)** and **(F)**: Same as panels (A) and (B), with distribution functions obtained from stochastic resonance models with period of 1470 years. From light to dark green the model parameters are:  $a=0.1, 0.2, 0.4$  and  $\sigma=0.38, 0.35, 0.27$  (see text), which generates on average 11 DO-events in 31 kyr. The important difference from the case shown in the panels above is that the Rayleigh's  $R$  and Std. dev. of residual in this case are calculated for the fixed period of 1470 yr. The red bars are ice-core data as above. The gray curves are the distributions for the exponential model repeated from the top panels. This shows that the SR model with  $a=0.1$  cannot be identified in a sample, since spurious coincidental periodicities will give a better match to the data than the 1470 yr cycle.

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